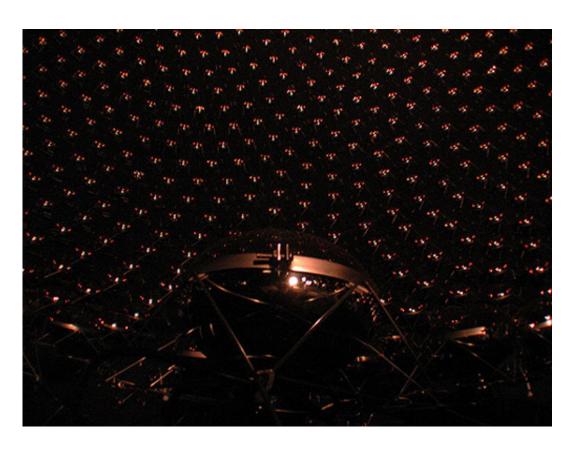
## **MiniBooNE**

Michel Sorel (Columbia University)

FNAL Users' Meeting, June 2003



- Physics
- Overview of the Experiment
- First look at the data
- Understanding the data

## MiniBooNE Timeline and Physics Potential



- •Beamline and detector completed
  - •MiniBooNE is SN-live
    - •First light from beam neutrinos
      - •Measure  $\nu_{\mu} \rightarrow \nu_{\mu}$
      - Measure  $\sigma(\nu_{\mu})$  ratios
      - •Measure strange spin of the nucleon
        - •Results on exotic searches
          - •Measure absolute  $\sigma(\nu_{\mu})$
          - •Measure  $\nu_{\mu} \rightarrow \nu_{e}$
          - •More  $\nu/\bar{\nu}$  running?
          - •Build 2nd detector (BooNE)?

#### The MiniBooNE Collaboration

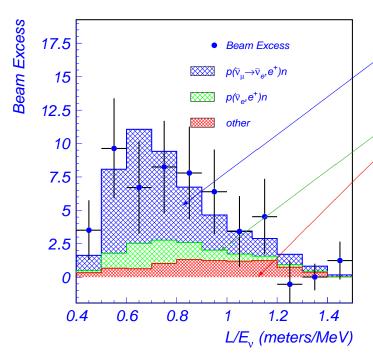
- ▶ Y. Liu, I. Stancu, University of Alabama
- ▶ S. Koutsoliotas, Bucknell University
- ▶ C. Green, University of California, Riverside
- E. Hawker, R. A. Johnson, J. L. Raaf, University of Cincinnati
- ▶ T. Hart, E. D. Zimmerman, University of Colorado
- L. Bugel, J. M. Conrad, J. Formaggio, J. M. Link, J. Monroe, M. H. Shaevitz, M. Sorel, G. P. Zeller, **Columbia University**
- D. Smith, Embry Riddle Aeronautical University
- L. Bartoszek, C. Bhat, S. J. Brice, B. C. Brown, D. A. Finley, B. T. Fleming, R. Ford, F. G. Garcia, P. Kasper, T. Kobilarcik. I. Kourbanis, A. Malensek, W. Marsh, P. Martin, F. Mills, C. Moore, P. Nienaber, E. Prebys, A. D. Russell, P. Spentzouris, R. Stefanski, T. Williams, Fermi National Accelerator Laboratory
- D. C. Cox, J. A. Green, H.-O. Meyer, R. Tayloe, Indiana University
- ▶ G. T. Garvey, W. C. Louis, G. McGregor, S. McKenney, G. B. Mills, E. Quealy, V. Sandberg, B. Sapp, R. Schirato, R. Van de Water, D. H. White, Los Alamos National Laboratory
- ▶ R. Imlay, W. Metcalf, M. Sung, M. O. Wascko, Louisiana State University
- J. Cao, Y. Liu, B. P. Roe, University of Michigan
- D. A. O. Bazarko, P. D. Meyers, R. Patterson, F. C. Shoemaker, H. Tanaka,

**Princeton University** 

- $\bullet \sim 60 \text{ scientists}$
- 13 institutions



#### The LSND Result



Oscillation signal expectation

### Backgrounds

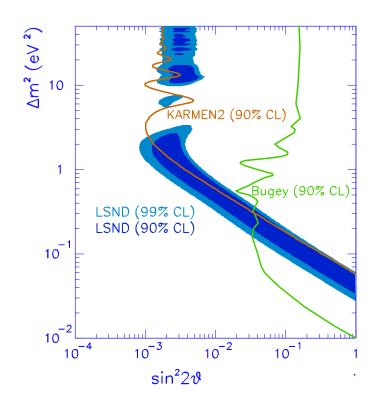
- $4\sigma$  excess of  $\bar{\nu}_e$  events in a  $\bar{\nu}_\mu$  beam
- Evidence for  $\bar{\nu}_{\mu} \to \bar{\nu}_{e}$  oscillations

• Two-neutrino oscillations:

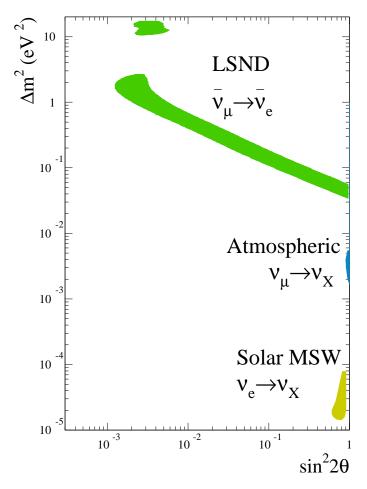
$$\begin{pmatrix} \nu_e \\ \nu_\mu \end{pmatrix} = \begin{pmatrix} \cos \vartheta & \sin \vartheta \\ -\sin \vartheta & \cos \vartheta \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix}$$
$$\Delta m^2 = m_2^2 - m_1^2$$

• Oscillation probability:

$$P_{\bar{\nu}\mu\to\bar{\nu}e} = \sin^2 2\theta \sin^2(1.27\Delta m^2 L/E)$$

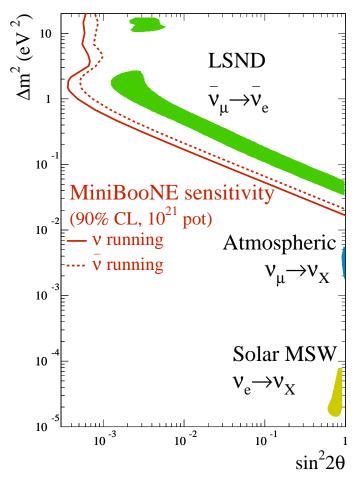


## Beyond minimal extensions of the SM?



- Three distinct neutrino oscillation signals, with:  $\Delta m_{sol}^2 + \Delta m_{atm}^2 \neq \Delta m_{LSND}^2$
- LEP: only three, light, weakly-interacting neutrinos
- Possible ways out...
  - 1. One experiment is not due to oscillations
  - 2. Active-to-sterile neutrino oscillations
  - 3. CPT violation

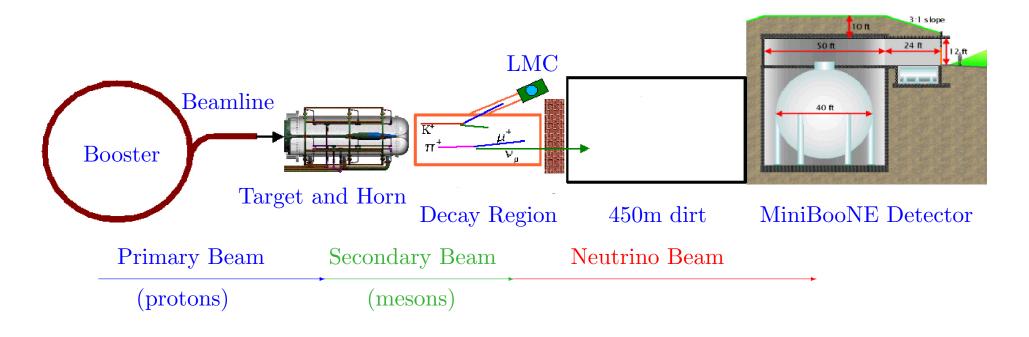
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- Possible ways out...
  - 1. One experiment is not due to oscillations
  - 2. Active-to-sterile neutrino oscillations
  - 3. CPT violation
- MiniBooNE can address all these possibilities:

- 1. Check LSND with different systematics, higher statistics, similar L/E
- 2.  $\nu_{\mu} \rightarrow \nu_{e} \text{ versus } \nu_{\mu} \rightarrow \nu_{\mu}$
- 3.  $\nu$  versus  $\bar{\nu}$  running mode

#### **Booster Neutrino Beam**



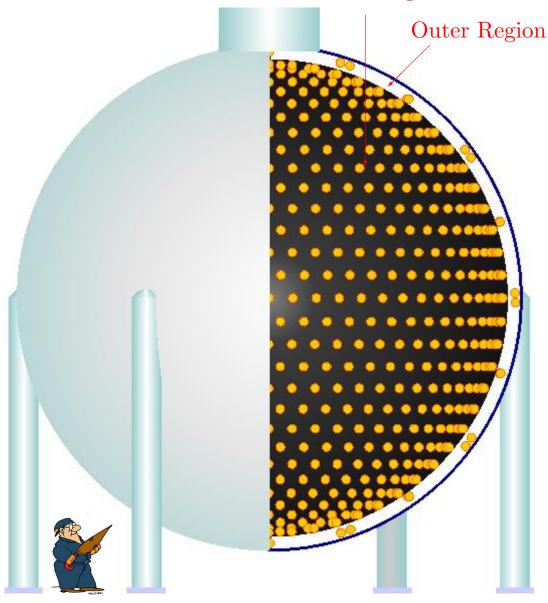
**Primary Beam:** 8 Gev protons from Booster,  $8 \cdot 10^{-6}$  duty factor

**Secondary Beam:** mesons are produced from protons striking Be target, focused by horn, and monitored by "Little Muon Counters" (LMC)

Neutrino Beam: neutrinos from meson decay in 50m pipe, pass through 450m of dirt (and oscillate?) to reach MiniBooNE detector

#### MiniBooNE Detector

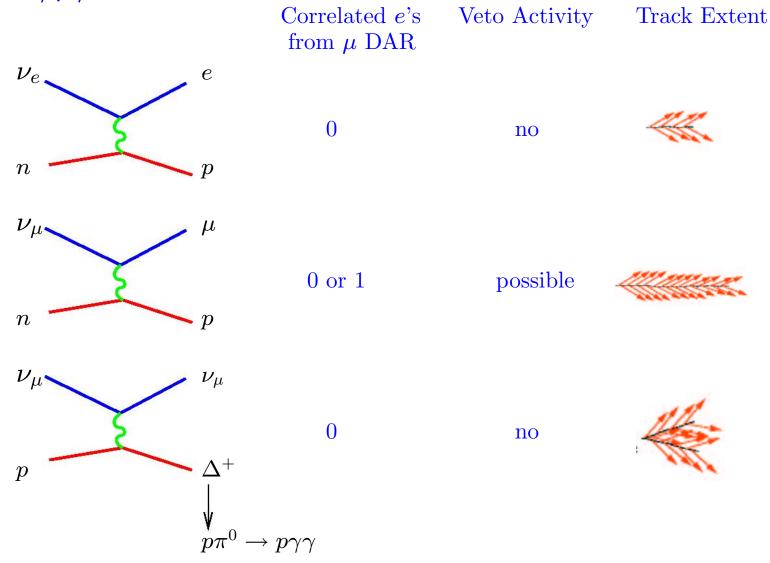
Inner Region



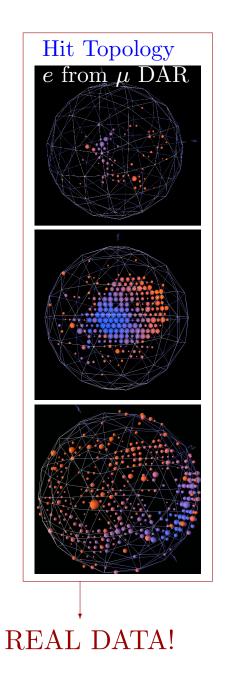
- 12m in diameter sphere filled with 800t of pure mineral oil
- Light tight inner region with 1280 8" PMTs (10% coverage)
- 240 PMTs in outer region (>99% veto efficiency)
- Neutrino interactions in oil produce:
- 1. Prompt, ring-distributed Cherenkov light
- 2. Delayed, isotropic scintillation light

#### Particle ID

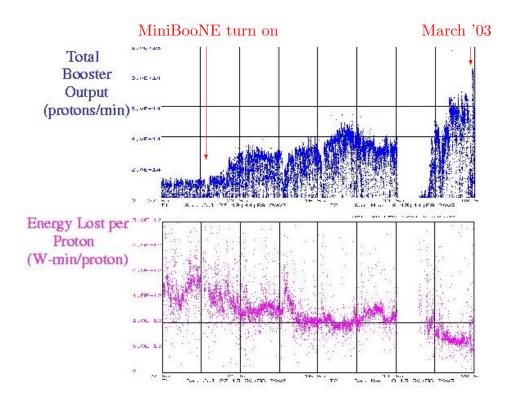
•  $e/\mu/\pi^0$  separation:



• Nuclear recoil: use scintillation/Cherenkov fraction

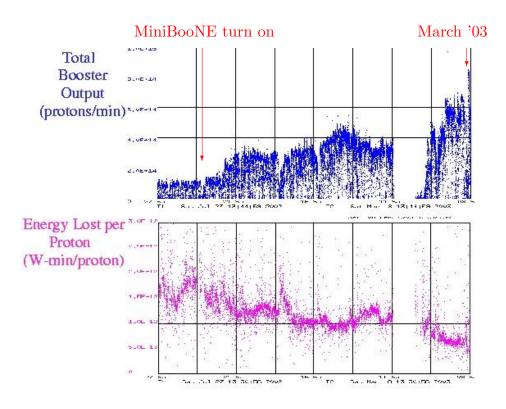


#### **Booster Performance**



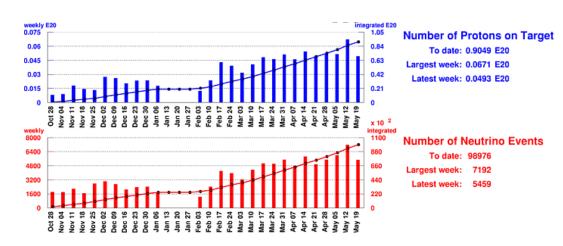
- Booster is working harder than ever!
- Steady increase of rate of delivered protons
- Currently factor of 2-3 below designed intensity. Designed intensity can be reached with planned Booster upgrades
- THANK YOU BEAMS DIVISION!
- Booster effort is already paying off...

#### **Booster Performance**



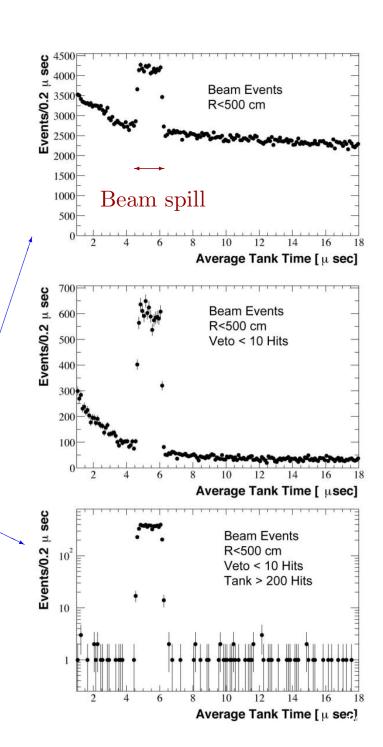
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- Booster effort is already paying off...

- Currently at 10% of 10<sup>21</sup> protons on target
- 100,000  $\nu_{\mu}$  event candidates collected so far



## **Beam Event Timing**

- Beam comes in spills at  $\sim 3$  Hz (we hope to bring this to 5Hz)
- Each spill: 82 bunches separated by 19ns  $\Rightarrow 1.6\mu s$  spill
- Trigger on signal from Booster; readout for  $19.2 \ \mu s$
- No high level analysis needed to see neutrino events over background!
- A few very simple cuts are sufficient to reduce the beam unrelated background to  $< 10^{-3}$
- We can even tell which RF bucket a neutrino comes from

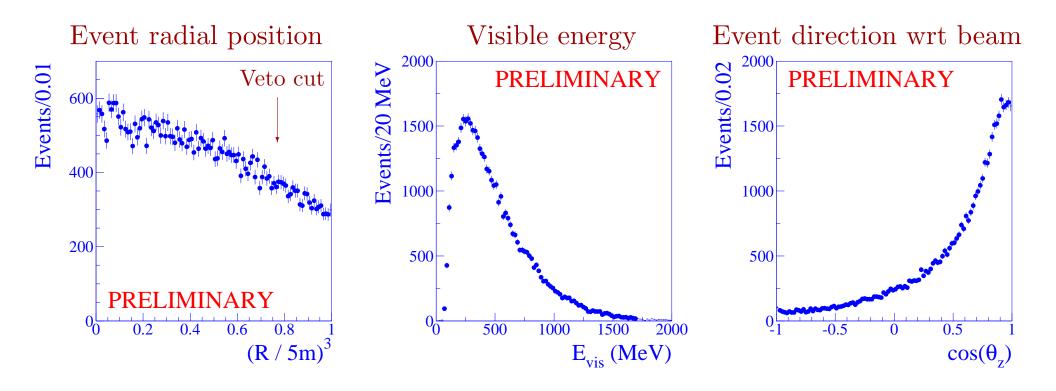


#### **Data Reconstruction**

- Select events with no particle ID requirements:
  - 1. In time with beam
  - 2. Center of event track within a 5m radius
  - 3. Contained event (low veto activity)
  - 4. Visible energy greater than endpoint for electron from  $\mu$  DAR

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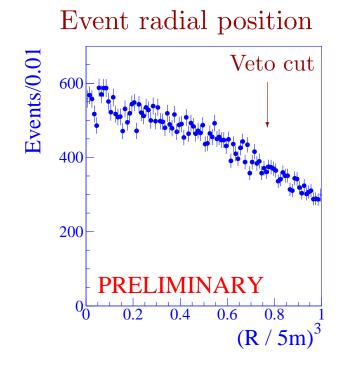
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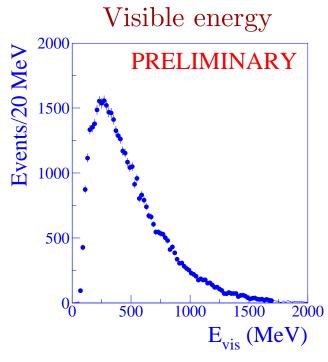


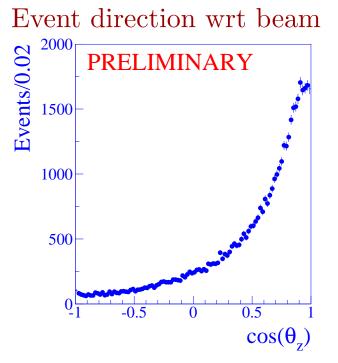
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- Reconstruction works well
- In the process of assigning uncertainties to MC expectations
- Data/MC comparisons in the fall, with first physics results





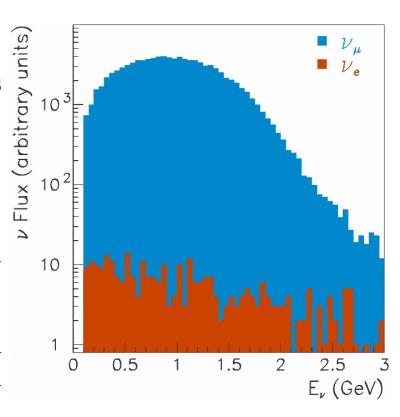


## Understanding the neutrino fluxes

- Most neutrinos from:  $\pi^+ \to \mu^+ \nu_\mu$
- For  $\nu_{\mu} \rightarrow \nu_{e}$  search, important background is intrinsic  $\nu_{e}$  background in the beam:

$$K \to \pi e \nu_e, \qquad \pi^+ \to \mu^+ \nu_\mu \\ \hookrightarrow \bar{\nu}_\mu e^+ \nu_e$$

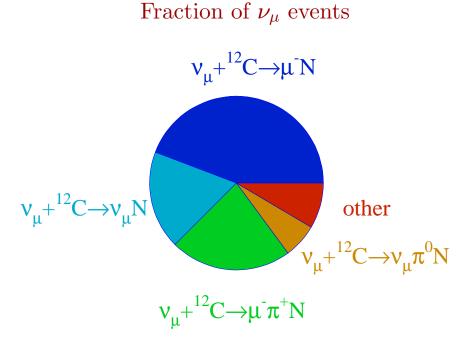
- Flux uncertainty dominated by uncertainty on  $\pi$ , K production in p-Be interactions
- Constrain neutrino flux predictions with existing  $\pi$  production data, BNL E910, and CERN HARP



- Internal cross-checks:
  - 1.  $\nu_e$  from  $K^+$  decays from high  $p_t$  muons in LMC (to be installed this summer)
  - 2.  $\nu_e$  from  $\mu^+$  decays  $\nu_\mu$  data and variable-length decay region

## Understanding the neutrino cross-sections

- At  $\sim$  1GeV, dominant processes are neutrino-nucleon quasi-elastic scattering and resonant  $\pi$  production
- Low energy regime and nuclear effects complicate things
- Example: final state interactions impact kinematics/rates and observed final states

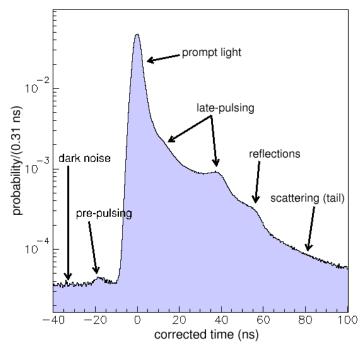


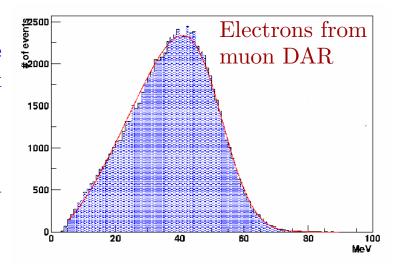
- MiniBooNE uses and develops the NUANCE  $\nu$  cross-section generator  $\Rightarrow$  world-wide collaboration within the neutrino physics community
- MiniBooNE will measure a variety of neutrino cross-sections
- Useful for other experiments as well (e.g. Super-K, K2K, MINOS)

## Understanding the detector

- Optical properties of the detector:
  - Dedicated off-situ measurements for measuring light production
  - Laser flask system measures light propagation and absorption in the detector
- Single PMT response:
  - Laser flask system measures the PMT charge and time response
- Track reconstruction:
  - Muon tracker plus scintillator cubes provide tracks with known direction, pathlength, vertex
- Energy scale and resolution:
  - Electron sample with known energy distribution from cosmic muons stopping in the detector

Timing distribution for laser events

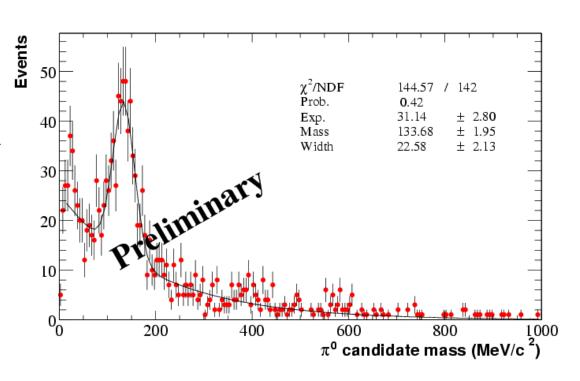




# Understanding the reconstruction, particle ID, event selection

- Build on the experience of previous Cherenkov detectors
- Extra handle of scintillation light
- Background to  $\nu_{\mu} \rightarrow \nu_{e}$  search: muons or  $\pi^{0}$ 's misidentified as electrons
- Goal: electron events selection with rejection at the level of  $10^3$  for muons,  $10^2$  for  $\pi^0$ 's

• Invariant mass from relatively pure  $\pi^0$  data sample is as expected:



## **Summary**

- Many accomplishments in first nine months of data taking
- Proton rate delivered by Booster has dramatically improved over time
- Further Booster upgrades are in the works to reach intended rate
- Detector works beautifully!
- Collected 100,000 neutrino event candidates so far
- Next step is to present in the fall:
  - 1. first physics results on  $\nu_{\mu}$  disappearance and cross-sections
  - 2. updated  $\nu_{\mu} \rightarrow \nu_{e}$  sensitivity
- Results on  $\nu_{\mu} \rightarrow \nu_{e}$  search to be expected in early 2005

#### **Extra-slides**

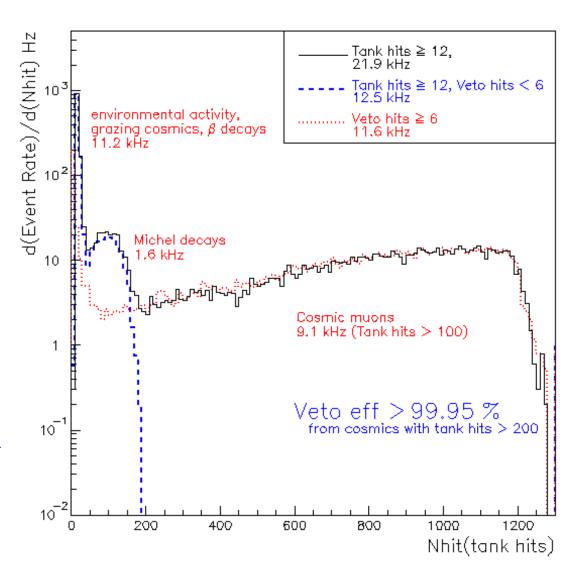
- MiniBooNE trigger
- $\nu_{\mu} \rightarrow \nu_{e}$  analysis
- $\nu_{\mu} \rightarrow \nu_{\mu}$  analysis

## MiniBooNE Trigger

• Typical trigger rates:

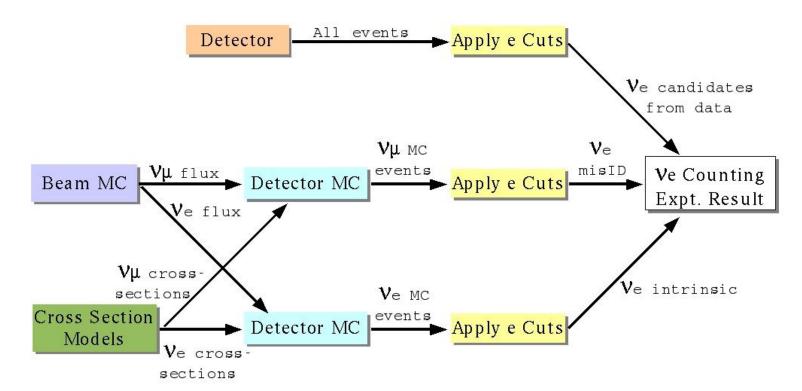
Trigger Type	Rate (Hz)
Beam	(currently) 3
Random	2
Laser flasks	1
Tank/Veto NHIT	1
Michel	1
Tracker/Cubes	1
Gamma/Beta	1
Supernova	11
Total	22

- Detector response understood down to a few MeV
- High veto efficiency



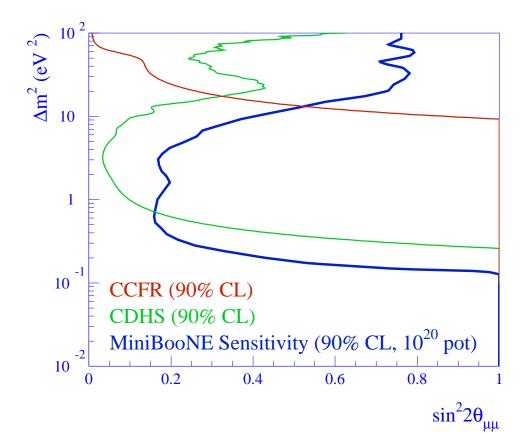
## $\nu_{\mu} \rightarrow \nu_{e}$ analysis scheme

- Ongoing efforts on all fronts needed for the analysis:
  - Neutrino fluxes
  - Neutrino cross-sections
  - Detector calibration
  - Particle ID, event selection algorithms



## $u_{\mu}$ disappearance analysis

- Compare predicted visible energy spectrum of  $\nu_{\mu}$  quasi-elastic events with data
  - search for a  $\nu_{\mu}$  disappearance signal
- Uncertainties in the flux and cross-section normalization are large
  - sensitivity comes from the energy distribution



(go back)